



1 **Top-of-the-atmosphere shortwave flux estimation from UV satellite observations: An** 2 **empirical approach using data from the A-train constellation**

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 4 Pawan Gupta^{1,2}, Joanna Joiner², Alexander Vasilkov^{3,2}, P. K. Bhartia²

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 6 [1] {Universities Space Research Association, Greenbelt, MD}
 7 [2] {NASA Goddard Space Flight Center, Greenbelt, MD}
 8 [3] {Science Systems and Applications, Inc., Greenbelt, MD}

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 10 Correspondence to: Pawan Gupta (pawan.gupta@nasa.gov)

11 12 **Abstract**

13 Estimates of top of the atmosphere (TOA) radiative flux are essential for the understanding of
 14 Earth's energy budget and climate system. Clouds, aerosols, water vapor, and ozone (O₃) are
 15 among the most important atmospheric agents impacting the Earth's short-wave (SW)
 16 radiation budget. There are several sensors in orbit that provide independent information
 17 related to these parameters. Having coincident information from these sensors is important for
 18 understanding their potential contributions. The A-train constellation of satellites provides a
 19 unique opportunity to analyze near-simultaneous data from several of these sensors. In this
 20 paper, retrievals of cloud/aerosols parameters and total column ozone (TCO) from the Aura
 21 Ozone Monitoring Instrument (OMI) have been collocated with the Aqua Clouds and Earth's
 22 Radiant Energy System (CERES) estimates of TOA SW flux (SWF). We use these data to
 23 develop a variety of neural networks that estimate TOA SWF globally over ocean and land
 24 using only OMI data as inputs. OMI-estimated TOA SWF reproduces the independent
 25 CERES data with high fidelity. The global mean daily TOA SWF calculated from OMI is
 26 consistently within ±1% of CERES throughout the year 2007. Application of our neural
 27 network to other ultraviolet sensors, both past and future, may provide unique estimates of
 28 TOA SWF. For example, the well-calibrated Total Ozone Mapping Spectrometer (TOMS)
 29 series could provide estimates of TOA SWF dating back to late 1978.



1 **1 Introduction**

2 The Earth's energy budget constrains the general circulation of the atmosphere and
3 determines the climate of the Earth-Atmosphere system; it is therefore also an indicator of
4 possible climate changes (Hatzianastassiou, et al., 2004). There is a long history of attempts to
5 estimate Earth's albedo and energy budget (Dines, 1917; Hartmann et al., 1986). With the
6 advent of the satellite remote sensing era, it became possible to directly measure the albedo of
7 the Earth. Subsequently, the shortwave energy balance at the TOA and the role of clouds,
8 aerosols, and trace gases has been studied using satellite measurements (Ramanathan et al.,
9 1989; Yu et al., 2006; Bellouin et al., 2005; Loeb et al., 2005; Patadia et al., 2008; Joiner et
10 al., 2009).

11 The Earth Radiation Budget Experiment (ERBE) was launched in October 1984 by the space
12 shuttle Challenger and provided long- and short-wave (SW) radiation parameter
13 measurements. Top-of-atmosphere short-wave (TOA SW) radiative parameter estimates from
14 ERBE (Barkstrom, 1984), Barkstrom and Smith (1986) showed that clouds approximately
15 double the albedo of Earth from an estimated clear-sky value of 0.15 to its average all-sky
16 value of 0.3 (Ramanathan et al., 1989; Harrison et al., 1990). The next generation of
17 broadband instruments, the Cloud and the Earth's Radiant Energy System (CERES), draws
18 heavily on ERBE heritage. Since its first launch in 1997 onboard the NASA Tropical Rainfall
19 Measurement Mission (TRMM), CERES has provided continuous observations that can be
20 used to understand the role of clouds and the energy cycle in global climate change (Wielicki
21 et al., 1995; Loeb et al., 2012).

22 Continuous and coincident measurements of radiative fluxes and atmospheric components
23 facilitate research studies to estimate and understand the role of different atmospheric
24 components on the planetary energy balance. Although CERES provides state-of-art estimates
25 of TOA radiative fluxes, it does not make measurements of individual atmospheric
26 components that impact those fluxes. Several studies have utilized aerosol and cloud
27 information from high spatial resolution MODerate resolution Imaging Spectroradiometer
28 (MODIS) measurements to quantify their impact on TOA fluxes (Yu et al., 2006; Patadia et
29 al., 2008; Zhang et al., 2005b; Loeb et al., 2005; Oreopoulos et al., 2009). Several-attempts
30 have also been made to convert narrowband radiances into broadband fluxes using regression
31 or more sophisticated statistical approaches (Chevallier et al., 1998; Hu et al., 2002;
32 Domenech and Wehr 2011; Vázquez-Navarro et al., 2012).



1 The Ozone Monitoring Instrument (OMI), flying on NASA's Aura satellite since 2004,
2 provides information about components important for the Earth's SW radiation budget
3 including the effective cloud/aerosol fraction (Stammes et al., 2008; Joiner and Vasilkov,
4 2006) and total column ozone (TCO) (Veefkind et al., 2006; McPeters et al., 2008; Kroon et
5 al., 2008). OMI-retrieved parameters can be utilized to understand their role in the Earth's SW
6 energy budget.

7 To model the spatial and temporal distribution of the TOA short-wave flux (SWF) requires a
8 description of the components that control the transfer of solar radiation within the Earth-
9 Atmosphere system. When required parameters are missing or incomplete, a statistical
10 approach is an alternative for estimation of TOA SWF. Here, we develop an artificial neural
11 network (NN) model to estimate TOA SWF. Artificial neural networks are algorithms that
12 simulate biological neural networks by learning and pattern recognition (Bishop, 1995). NNs
13 have been used by many scientific disciplines, including Earth science, to identify patterns
14 and extract trends in imprecise and complicated non-linear data (e.g., Lee et al., 1990; Gupta
15 and Christopher, 2009). In radiation studies, NNs have been used to estimate TOA and
16 surface short-wave fluxes based on radiative transfer calculations with or without data from
17 satellites (e.g., Krasnopolsky et al., 2008, 2010; Takenaka et al., 2011; Vázquez-Navarro et
18 al., 2012; Jiang et al., 2014). CERES TOA flux algorithms have also used NNs to generate
19 Angular Distribution Models (ADMs) in the absence of sufficient high-resolution imager
20 information for reliable scene identification (Loukachine and Loeb, 2003; 2004).

21 In this study, we utilize OMI cloud, and ozone products along with other ancillary data to
22 estimate TOA SWF. We develop NNs that take OMI-derived quantities as inputs and provide
23 CERES-equivalent TOA SWF as the output. The trained NN models are optimized to run
24 with data sets from OMI or similar sensors and can be applied generally to different seasons
25 and years. For example, the neural network-based models we develop here can be applied to
26 similar measurements from the Total Ozone Mapping Spectrometer (TOMS) instruments. The
27 main objective of this study is to assess how well TOA SWF can be estimated using OMI
28 cloud and ozone products with NNs when CERES data are used for training. The developed
29 NNs can then be applied to other instruments with similar accuracy.

30 The paper is organized as follows: Section 2 describes the various satellite data sets utilized in
31 the study. Section 3 discusses the development of NN models including the selection of input



1 parameters. Section 4 evaluates our NN estimation of TOA SWF using independent CERES
2 data over ocean and land. Section 5 summarizes the results and discusses future work.

3 2 Satellite Data Sets and Coincident Sampling

4 Under clear-sky conditions, TOA SWF is affected by the Earth's surface properties,
5 atmospheric absorbers such as water vapor, ozone, and aerosols, and scattering by air
6 molecules and aerosols particles. Over ocean, surface properties can be characterized by
7 ocean color and roughness of the ocean surface. Under cloudy sky conditions, cloud optical
8 properties such as the cloud optical thickness, geometrical cloud fraction, effective radius, and
9 phase function affect TOA SWF. In clear and cloudy skies, the solar zenith angle (SZA) and
10 Sun-Earth distance (SED) impact the TOA SWF.

11 In this work, we make use of data sets mainly from two passive sensors in A-train
12 constellation of satellites that fly within 15 minutes of each other: 1) Aura OMI with an
13 equatorial crossing time of $\sim 13:45 \pm 15$ minutes local time and 2) Aqua CERES with an
14 equatorial crossing time of $\sim 13:30$ local time. We primarily use 2007 data over global oceans
15 for the training, testing, and validation of neural network models. Starting around 2008, OMI
16 experienced an anomaly presumably due to material outside the sensor that adversely affects
17 the quality of the level 1B and level 2 data products in a portion of its 60 rows across the
18 swath. Our study focuses on data in 2007 that are not significantly affected by these
19 anomalies.

20 2.1.1 CERES

21 The first CERES instrument flew on the TRMM satellite, launched in November 1997, and
22 provided data until 2000. Five CERES instruments are currently operating; two on NASA's
23 Terra satellite (FM1 and FM2), two on NASA's Aqua satellite (FM3 and FM4), and one on
24 the Suomi National Polar-orbiting Partnership (NPP) satellite (FM5). These CERES
25 instruments provide radiometric measurements of the Earth's atmosphere from three
26 broadband channels: 1) A shortwave channel to measure reflected sun light ($0.3\text{--}5\text{ }\mu\text{m}$), 2) a
27 long wave channel to measure Earth-emitted thermal radiation in the window region ($8\text{--}12\text{ }\mu\text{m}$), and 3) a total channel to measure radiation from 0.3 to $200\text{ }\mu\text{m}$.

29 CERES radiances are converted to TOA fluxes using angular dependent models (ADMs). The
30 CERES science team has an extensive database of ADM's for clear- and cloudy-sky over both
31 land and ocean (Loeb et al., 2005). The ADMs heavily depend upon the observed scene type



1 and are sensitive to surface characteristics as well as cloud and aerosol optical properties
2 (Loeb et al., 2003; Zhang et al., 2005a; Patadia et al., 2011). The ADMs over ocean are
3 dependent upon wind speed and aerosol optical thickness along with sun-satellite geometry
4 (Zhang et al., 2005a).

5 The Aqua spacecraft carries two identical CERES instruments: one operates in a cross-track
6 scan mode (FM3) and the other in a biaxial scan mode (FM4). Measurements from the biaxial
7 scan mode were used to develop the ADMs; this provided considerable improvement over the
8 previous generation of instruments including the ERBE (Loeb et al., 2003; 2007).

9 This study uses the Single Scanner Footprint (SSF, Edition 3A) TOA SWF obtained from the
10 Aqua CERES FM3. The SSF product is a merge of CERES parameters with coincident cloud
11 and aerosol parameters derived from the Aqua MODIS (Loeb et al., 2003). The high-
12 resolution ($1 \times 1 \text{ km}^2$ at nadir) MODIS imager data are used to characterize the clear and
13 cloudy portions of the larger CERES pixel ($20 \times 20 \text{ km}^2$ at nadir).

14 2.2 OMI

15 OMI provides hyper-spectral measurements of Earth-backscattered sunlight from UV to
16 visible wavelengths ($\sim 270\text{-}500 \text{ nm}$) with a spectral resolution of the order of 0.5 nm (Levelt et
17 al., 2006). Its spatial resolution is $13 \times 24 \text{ km}^2$ at nadir with a swath width of about 2600 km .
18 Cloud, aerosol, and total column ozone (TCO) products from OMI are used in this study.
19 Specifically, the cloud-aerosol Optical Centroid Pressure (OCP), effective cloud fraction (f_c),
20 Lambertian-Equivalent Reflectivity (LER) at 354.1 nm , Solar Zenith Angle (SZA), Relative
21 Azimuth Angle (RAA), and Viewing Zenith Angle (VZA) are obtained from the OMI cloud
22 products as detailed below, and Aerosol Index (AI) and TCO are obtained from the OMI-
23 TOMS total ozone product (OMTO3, version 8.5, collection 3) (McPeters et al., 2008).

24 Cloud-aerosol OCP, also known as effective cloud pressure, is a measure of the reflectance-
25 weighted pressure reached by incoming solar photons (Joiner et al., 2012). It is distinct from
26 the cloud-top pressure (CTP). While CTP is the more important parameter needed for TOA
27 long-wave flux, OCP is more related to atmospheric absorption in the short-wave. OCP is
28 derived from OMI observations using two different methods (Stammes et al., 2008): 1)
29 filling-in of solar Fraunhofer lines from rotational-Raman (RR) scattering in the UV (the
30 OMCLDRR product) (Joiner and Bhartia, 1995; Joiner et al., 2004) and 2) collision-induced
31 oxygen absorption ($\text{O}_2\text{-O}_2$) at 477 nm (the OMCLDO2 product) (Stammes et al., 2008;



1 Acarreta et al., 2004). Unless otherwise specified, we use the f_c and OCP from OMCLDRR
 2 product here.

3 OMI cloud and trace-gas algorithms use a simplified mixed Lambertian cloud model to
 4 estimate observed radiances I_m . In this scenario, a pixel is modeled as having components
 5 from clear and cloudy sub-pixel weighted using an effective cloud fraction f_c , i.e.,

$$I_m = I_g (1 - f_c) + I_c f_c \quad (1)$$

6 where I_g and I_c are the radiances computed in the Rayleigh atmosphere for Lambertian
 7 surfaces corresponding to the clear and cloudy portions of the scene, respectively; f_c is defined
 8 as the fraction of the Lambertian cloud covering the pixel and is related to both the geometric
 9 cloud fraction and cloud optical thickness. It contains information similar to the Lambertian-
 10 equivalent reflectivity of the scene (related to cloud and surface reflectivities). However,
 11 because it attempts to account for variations in the Earth's surface reflectivity, it is a more
 12 spectrally invariant quantity and therefore potentially more highly correlated with TOA SWF.

13 Formally, the effective cloud fraction is wavelength-dependent because it is defined by
 14 spectral quantities (Stammes et al., 2008). We conducted a simulation experiment to evaluate
 15 the wavelength dependence of f_c . In this experiment, we simulate observed TOA radiances as
 16 a weighted sum of the clear-sky and cloudy radiances, i.e.,

$$I_m = I_g (1 - f_g) + I_c f_g \quad (2)$$

17 where I_c is the cloudy radiance computed with a plane-parallel cloud model that depends on
 18 cloud optical thickness, and f_g is the geometrical cloud fraction. In our simulation, clouds
 19 have a vertically uniform distribution of the extinction coefficient and phase function. We use
 20 a cloud top height of 5 km and a cloud layer thickness of 1 km. The assumed cloud optical
 21 depth of 20 is spectrally-independent within the 320-1400 nm wavelength range. The
 22 spectrally-independent optical thickness is a good approximation for clouds with sufficiently
 23 large particles (Deirmendjian, 1969). We neglect gaseous absorption in the specified spectral
 24 range. Three models of cloud phase function are used: 1) ice crystals with an effective
 25 diameter of 60 μm (Baum et al., 2014), 2) C1 water droplets with an effective diameter of 12
 26 μm (Deirmendjian, 1969), and 3) the Heneye-Greenstein (HG) model (e.g. van de Hulst and
 27 Irvine, 1963) with an asymmetry parameter of 0.85. We use a simplified model of the spectral
 28 ground reflectance: $R_g = 0.05$ at $\lambda < 700 \mu\text{m}$, $R_g = 0.2$ at $\lambda > 700 \mu\text{m}$. We then calculate f_c by



1 inverting Eq. 1 assuming a Lambertian cloud with a reflectivity of 0.8; this is commonly used
 2 for trace-gas algorithms (Stammes et al., 2008).

3 Figure 1 shows the spectral dependence of f_c calculated for $f_g = 0.5$, $SZA = 45^\circ$, and at nadir
 4 for the three phase functions assuming a cloud single scattering albedo of unity. It can be seen
 5 that f_c is nearly invariant with wavelength over a wide spectral range; it changes by only a few
 6 percent even for the steep change in the ground reflectance (simulating the so-called red-edge)
 7 at 700 nm. This result holds when other input parameters in our simulation are varied. The
 8 near spectral invariance of f_c suggests that it will be highly correlated with TOA SWF and
 9 thus a good predictor in a statistical model of TOA SWF.

10 We have used the following modified cloud fraction parameter, f_{c_mod} , as a predictor to
 11 estimate TOA short wave flux:

$$12 \quad f_{c_mod} = f_c \times \cos(SZA) \times \left(1/SED^2\right)$$

13 where SED is the sun-Earth distance. The modification accounts for variation in the incoming
 14 solar irradiance. If SED is excluded from the input parameters, this creates time-dependent
 15 biases in the estimated TOA SWF. Figure 2 demonstrates that the relationship between TOA
 16 SWF and f_{c_mod} is highly linear and that this single parameter captures much of the variability
 17 in TOA SWF.

18 **2.3 Ancillary Data**

19 In addition to OMI data, a SeaWiFs-derived chlorophyll concentration (Chl) climatology is
 20 used as an input predictor when $f_c < 1$. The Precipitable Water (PW) and 2 m surface wind
 21 speed (Wind) are also used as predictors; these are provided in the CERES SSF data set and
 22 are taken from the GEOS 4 reanalysis (Bloom et al., 2005).

23 **2.4 Coincident Sampling of OMI and CERES**

24 Because the sizes of the OMI (13 km x 24 km) and CERES (20 km²) pixels are similar at
 25 nadir, we perform a simple spatial collocation by finding the closest CERES pixel
 26 corresponding to each OMI pixel. OMI and CERES collocated pixels are only included in our
 27 training and validation samples when the distance between centers of OMI and CERES pixels
 28 is less than 20 km. We examine the frequency distribution of the distance between OMI and
 29 CERES pixels of all the collocated data sets and found that most of the collocated data (98%
 30 and 60%) have distances less than 20km and 10km, respectively. We do not include pixels



1 with viewing zenith angles $> 60^\circ$. At these angles, OMI and CERES pixels become
 2 significantly larger (~ 150 km for OMI and ~ 200 km for CERES in the cross track direction)
 3 and may contain different scene types. We also mask OMI pixels with $AI > 1$ to avoid heavy
 4 absorbing aerosol loaded scenes where the f_c and OCP are known to contain errors (Vasilkov
 5 et al., 2008). The quality-controlled collocated data are averaged on equal latitude and
 6 longitude grids of $1^\circ \times 1^\circ$ for training, testing, and validation of the neural networks.

7 Although the NN training includes data from CERES and other ancillary data sets but the
 8 trained NN provides TOA SWF similar to CERES using predominantly retrievals from OMI
 9 measurements. Therefore, the NN produced TOA SWF flux will be referred as OMI estimated
 10 SWF throughout the manuscript.

11

12 **3 Artificial Neural Network Model**

13 **3.1 General NN architecture and training approach**

14 The general neural network architecture has three layers of neurons: an input layer, a hidden
 15 layer, and an output layer with standard multi-layer network architecture. We use the same
 16 number of neurons in the hidden layer as in the input layer as this produced generally good
 17 result. The input layer has an identity activation function; all other layers are connected by a
 18 sigmoid activation functions (Equation 3).

$$y(x) = \frac{1}{1 + e^{-x}} \quad (3)$$

19 The network normalizes both input and output data sets with a unique linear mapping for each
 20 input and output parameter. Figure 3 provides an example of a schematic of the network used
 21 in our study. Here we used two different NN models: one with nine nodes or parameters
 22 (NNM1) and a second with seven nodes (NNM2) in the input layer. Both of these models
 23 have one node (TOA SWF) in the output layer. Figure 3 also lists the input parameters
 24 corresponding to the NNM1 and NNM2 models. The NNM1 model is optimized for ocean
 25 cases where the OMI $f_c < 1.0$, whereas NNM2 is optimized for cases where $f_c = 1$ (saturated
 26 cases).

27 Neural network-based models require optimized training to produce accurate outputs. Here we
 28 use a standard back propagation training algorithm (Hertz et al., 1991), where inputs are
 29 iteratively sent to the neural network. In back propagation, the hidden layer weights



1 associated with each input parameter are modified through the training process that minimizes
 2 errors between the targets and outputs (Bishop, 1995; Gardner and Dorling, 1998). After each
 3 iteration, the error is propagated backward through the network and weights are modified to
 4 bring the actual response of the network closer to the desired output in a statistical sense. The
 5 function minimized during the training is a sum of squared errors of each output for each
 6 training pattern. Once the network is trained, it can be evaluated using independent data (i.e.,
 7 not used in the training data set).

8 **3.2 Impact of different input parameters**

9 Here we examine the impact of using various input parameters on the derived neural
 10 networks. This exercise is performed using data with $f_c < 1$ with one month of the data over
 11 ocean (January 2007). Table 1 presents the performance of eight different NNMs, denoted
 12 models a through h, with various input parameters listed in Table 1 and described in more
 13 detail in sections 3.2.1-3.2.2.

14 **3.2.1 Inclusion of OMI UV-derived parameters**

15 In model a, we have combined the effects of SZA, SED and f_c into a single input parameter
 16 called f_{c_mod} , which is defined in the section 2.2. Use of this modified input parameter alone
 17 explains about 94% ($R=0.97$) of the variability in TOA SWF. As we add other parameters in
 18 models b to h, we observe small improvements in the OMI-estimated TOA SWF. Figure 4
 19 shows the spatial distribution of monthly mean OMI-CERES SWF differences for these
 20 models.

21 The TOA SWF is estimated from a measured radiance and therefore the observational
 22 geometry factors in. The addition of satellite viewing geometry parameters (VZA, RAA) to
 23 model a provides improvements in areas of high biases and reduces the standard deviation
 24 from 37.1 Wm^{-2} to 31.4 Wm^{-2} . Model c tests the ability of 354 nm reflectivity (LER) to
 25 predict TOA SWF in place of f_{c_mod} . Although the statistical parameters in table 1
 26 corresponding to models b and c are very similar, we note spatial differences in the OMI-
 27 CERES TOA SWF in Figure 4. Further analysis reveals that the f_c -based model b provides
 28 more accurate flux estimation as compared with the LER-based model for a larger range of
 29 fluxes.

30 The inclusion of TCO (model d) as an input parameter positively impacts TOA SWF
 31 estimation as shown in Figure 4 (d); the high positive biases in the tropical Pacific and Indian



1 oceans and in the region near 60°N have been reduced. The percentage of monthly mean
 2 OMI-CERES data that falls within $\pm 8\%$ increases from 91% in model c to 94% in model d.
 3 Model e adds cloud OCP to the input parameters included in model b. OCP also improves
 4 SWF estimates; the regions where improvement occurs are different from those improved by
 5 using TCO. Model f shows that when TCO and OCP are used together as input parameters,
 6 there is further improvement in SWF estimation. Although the global statistics in table 1 do
 7 not clearly reflect this improvement, Figure 4(f) shows that inclusion of OCP and O3 reduces
 8 biases in many regions, most prominently in the tropics. The percentage of total OMI samples
 9 (monthly mean) within $\pm 8\%$ of CERES increases from 92% in model b to 95% in model f.
 10 Apart from these parameters, we also evaluated the inclusion of AI as an input parameter (not
 11 shown here). We found that overall it does not significantly improve the results; however it
 12 does provide some improvement in regions with high AI values.

13 3.2.2 Addition of meteorological and other ancillary data

14 The impact of surface winds and total column water vapor (model g in Figure 4g) is more
 15 prominent in the tropics than in other regions. Inclusion of chlorophyll and LERs in model h
 16 removes some of the notable low biases in TOA SWF near the coast of Northern China,
 17 Caspian Sea, and Black Sea. Furthermore, model h corrects for negative biases in areas with
 18 high TOA SWFs, most likely due to the inclusion of LER. The model h produces 89% (99%)
 19 of OMI-estimated monthly mean TOA SWFs within $\pm 5\%$ ($\pm 12\%$) of CERES and is best of
 20 the eight models.

21 3.3 Consistency over time

22 We next examine the performance of the NN model h with respect to different input samples.
 23 Figure 5(a-b) presents the results from two different years over ocean. We first examine the
 24 robustness of the NN for detection of inter-annual variability. In this exercise, we trained with
 25 data from the first 15 days of January 2007 as above, and applied it to data from the entire
 26 months of January 2007 (Figure 5a) and January 2006 (Figure 5b). Figure 5a and b shows that
 27 the NN performance is consistent between years. Although the number of samples in January
 28 of 2006 and 2007 is a bit different, the NN model produces similar statistics. The color of
 29 each coincident pair ($10 \times 10 \text{ Wm}^{-2}$ intervals) represents the density (%) of the matchup. The
 30 solid black 1 to 1 line is shown with three dotted lines on both sides that represent envelopes
 31 of $\pm 5\%$, $\pm 10\%$, and $\pm 15\%$ OMI-CERES differences.



1 In the next test, we applied the same model (trained on January 2007) to July 2007. In this
 2 case, results were degraded as compared with application to January data. We then trained the
 3 same network (with the same input parameters) using a subset of data from July 2007 and
 4 applied it to the entire month of July 2007. Results were similar to those of training and
 5 application to January. This exercise suggests that we may need to use different models for
 6 different months or expand our training data set for application to different months.

7 We next use data from the 1st day of each month of 2007 for training and data from the 16th
 8 day of each month of 2007 for evaluation and gridded the data at 1° latitude by 1° longitude
 9 resolution. The comparisons with CERES using the training and validation data are consistent
 10 as shown in Figure 6. The mean bias in both training and validation data sets is close to zero
 11 whereas standard deviation remains stable and close to 30 Wm⁻² in two independent model
 12 runs. The almost identical values of statistical parameters for training and validation data
 13 demonstrate that the neural network has been well trained. For example, there is a high degree
 14 of linear correlation ($R = 0.98$ or $R^2 = 0.96$) and slopes close to 1 (0.96) in both training and
 15 validation comparisons. Further analysis shows that 83%, 70% and 43% of TOA SWF
 16 estimated from OMI (training and validation data combine) lie within the 15%, 10%, and 5%
 17 of the CERES TOA SWF, respectively. The global standard deviation of the daily OMI-
 18 CERES is about 30 Wm⁻².

19 Further evaluation of the entire year reveals that this NN is appropriate for all months.
 20 Therefore this model will be used for subsequent analysis in this study. Creating more
 21 networks as a function of scene type or for different latitude belts or even for different
 22 months/seasons will improve results in certain regions. However, based on our results, we
 23 simplified the approach by minimizing the number of networks.

24 **3.4 Case of $f_c = 1$**

25 About 1-2 % of total coincident data correspond to $f_c = 1$, typical of overcast conditions with
 26 optically thick clouds. These cases were modeled using a simpler neural network model with
 27 inputs of LER, SZA, VZA, RAA, OCP, O3, and PW, the surface-related parameters (surface
 28 wind speed and chlorophyll content) do not produce a significant impact for ECF=1 and have
 29 therefore been removed. Subsequent results use combined output from the two separate
 30 models for $f_c < 1$ and $f_c = 1$.

31



1 **4 Results and Discussion**

2 **4.1 Bias and RMSE as a function of effective cloud fraction**

3 Figure 7 presents the Root Mean Square Error (RMSE), RMSE Normalized by CERES flux
 4 (NRMSE in %), data sample (%), and bias (%) for 5% ECF bins. This analysis includes both
 5 training and validation data as presented in Figure 6. The RMSE varies between about 24-35
 6 Wm^{-2} and continuously increases with cloud fraction (and observed flux). The NRMSE, on
 7 the other hand, continuously decreases with ECF from about 18% for 5% ECF to ~6% for
 8 overcast conditions. The bias represents the mean error (in %) for each ECF bin. The mean
 9 global bias shows more variability than RMSE and is highest (2.9%) for about 10% ECF. The
 10 bias decreases sharply from 2.9% at $f_c = 0.1$ to about 1.2% at $f_c = 0.4$. The bias remains low
 11 ($< 1.2\%$) for $f_c > 0.4$ (usually associated with frontal or deep convective clouds). The higher
 12 biases for lower f_c (usually associated with thin cirrus and broken clouds) are likely related to
 13 higher noise and uncertainties in OMI cloud parameters. For example, (Joiner et al., 2012)
 14 showed that cloud OCP errors increase with decreasing f_c . The biases may also be related to
 15 absorbing aerosol in the scene, particularly when it overlies clouds. This will be illustrated in
 16 more detail below as we show spatial variations in OMI-CERES differences.

17 **4.2 Effects of Spatial and Temporal Averaging**

18 In order to evaluate the NN performance at different spatial and temporal scales similar to
 19 those used by the climate community, we use data from July 2007. Figure 8 presents a
 20 comparison of daily CERES and OMI TOA SWF over ocean for 6 spatial scales: the OMI
 21 native pixel ($13 \times 24 \text{ km}^2$ at nadir) and 0.5° , 1° , 2° , 5° , and 10° gridded spatial resolutions.
 22 Statistical parameters for these comparisons are reported in Table 2. As expected, the pixel
 23 level data are much noisier than the gridded data owing to collocation noise in partly cloudy
 24 cases, but the slope (0.96) is still close to 1, and the linear correlation coefficient is 0.96 with a
 25 standard deviation of 47.7 Wm^{-2} . Below 300 Wm^{-2} , where the sample density is highest, the
 26 NN slightly underestimates the CERES SWF. The mean bias of the OMI-estimated SWF with
 27 respect to CERES is -1.4 Wm^{-2} . This bias may be due to a combination of effects including
 28 uncertainties in the input parameters as well as the limitations of the NN model itself. For
 29 example, we have excluded pixels with a clear signature of absorbing aerosols (OMI derived
 30 UV aerosol index > 1) where OMI effective cloud fractions and pressures may be in error in
 31 both the training and validation data. However, in some regions where smoke and dust
 32 overlaying clouds is common (e.g., west coast of Africa), pixels with erroneous cloud



1 fractions owing to small amounts of absorbing aerosol may be present in both the training and
 2 validation data. This may produce errors in the NN model and will be examined in more detail
 3 below.

4 For the daily data, as the spatial averaging scales increase from 0.5° to 10° , the OMI-estimated
 5 SWF becomes almost identical to CERES; the correlation coefficient increases from 0.97 to
 6 0.99, and the slope increases from 0.96 to 0.97. The percentage of OMI data that falls within
 7 $\pm 5\%$ of CERES increases from 37% for 0.5° to 69% for 10° grids. About 87% percent of
 8 OMI-estimated 2° gridded daily mean TOA SWFs are within 15% of CERES data.

9 Figure 9(a-d) shows 2D histograms of monthly mean gridded data over ocean at 0.5° , 1° , 2° ,
 10 and 5° spatial resolutions, respectively. The monthly inter-comparisons of OMI and CERES
 11 SWF show excellent agreement at all spatial resolutions with correlation coefficients of 0.99
 12 and slopes of 0.98 (Table 2). The global mean biases vary between -1.8 and 0.25 Wm^{-2} . The
 13 standard deviations vary between 6.6 and 12.9 Wm^{-2} for the different spatial resolutions.
 14 Ninety seven percent of monthly mean 1° OMI estimated TOA SWFs are within 15% of those
 15 derived from CERES, and 93% are within 10%.

16 4.3 Spatial Distribution of TOA SWFs over ocean

17 Figure 10 presents the spatial distribution of 1° monthly mean (July 2007) TOA SWF from
 18 CERES (Fig. 10a) and the difference with the OMI in Wm^{-2} (Fig. 10b) and percent difference
 19 (Fig. 10c). There are subtle differences between the NN and CERES estimates of TOA SWF
 20 as shown in Figure 10b and c. The OMI-CERES histograms (Fig. 10d) show that for 44%
 21 (79%) samples, NN fluxes are within ± 2 (± 5) % of CERES fluxes. About 9% of the samples
 22 have biases of $\pm 8\%$ or more. Overall, the northern hemisphere shows better agreement than
 23 southern hemisphere during July (boreal summer). This could be due to larger errors in the
 24 OMI cloud products at higher solar zenith angles. The low biases on the west coast of Africa
 25 may be due to the presence of absorbing aerosols, particularly when they occur over clouds.
 26 The striped pattern in the southern hemisphere (latitudes $>40^\circ\text{S}$) is mainly associated with high
 27 viewing zenith angles in conjunction with high solar zenith angles that occur on one side of
 28 the swath.

29 Figure 11 similarly shows differences between CERES and OMI TOA SWF over ocean
 30 derived using f_c and OCP from the OMI O2-O2 product in place of the OMI RRS product.
 31 Because there are slight differences in the two cloud products, we retrained the network with



1 OMI O₂-O₂ cloud parameters to be consistent. The use of the O₂-O₂ f_c and OCP improves
2 the accuracy of the estimated TOA SWF. The regions of improvement include the west coasts
3 of South America and southern Africa and some parts of Indian Ocean. The O₂-O₂ cloud
4 product, which uses visible wavelengths, is less affected by absorbing aerosol; this may
5 explain the improvement in these areas where absorbing aerosol, especially over clouds, is
6 common. However, negative biases remain over large regions off the west coast of Africa.

7 Figure 12a shows a time series of daily global mean values of TOA SWF over ocean from
8 OMI and CERES for 2007. Both instruments show almost identical daily variations with
9 differences within $\pm 1\%$. Figure 12b, c provides monthly averaged (July 2007) CERES and
10 OMI zonal and meridional means of TOA SWF. The OMI-derived TOA SWF is able to well
11 reproduce the variability shown in the CERES data.

12 **4.4 Spatial Distribution of TOA SWFs over Land**

13 We developed a land NN model that utilizes most of the input parameters from our ocean NN
14 (using OMI RRS cloud parameters). For surface characterization, we use a monthly
15 climatology of surface broad-band albedo in place of the chlorophyll concentration and
16 surface wind speed. The albedo product is derived using a combination of CERES and
17 MODIS observations at 1 degree spatial resolution (Rutan et al., 2009).

18 Figure 13 shows results from the OMI-derived CERES-trained NN that produces TOA SWF
19 over land. Statistical comparison with CERES over land provides results similar to those over
20 ocean. The NN performs well over Asia and parts of Europe and the Americas. The OMI-
21 based NN tends to underestimate TOA SWF over the high albedo desert areas of Northern
22 Africa, Australia, and also over some regions of South America. Note large differences that
23 occur in coastal regions may be due to imperfect collocations.

24 **5 Summary and Conclusions**

25 We have developed a neural network approach to estimate TOA SWF based primarily on UV
26 measurements using the Aura OMI with Aqua CERES data used for training. One year of data
27 from OMI and CERES has been used to train/validate/analyze several separate neural
28 networks for different conditions, which together provide estimation of TOA SWF under all-
29 sky conditions. The most important input parameters are effective cloud fraction and sun-
30 satellite geometry. Total column ozone and cloud optical centroid pressure from OMI, as well
31 as surface-related parameters, provide secondary positive impacts.



1 Independent validation at different spatial and temporal scales shows that the OMI NN-based
2 approach reproduces CERES-derived TOA SWF with high fidelity. Correlation coefficients
3 for all comparison are > 0.95 , and slopes are close to unity. A high percentage of OMI-
4 estimated monthly mean TOA SWF at 0.5° spatial resolution over global oceans (97%) falls
5 within 15% of CERES. The global mean bias in pixel level data of about -1.4 Wm^{-2} over
6 oceans with respect to CERES is likely due in part to errors in OMI cloud parameters that
7 occur in the presence of absorbing aerosols.

8 We plan to apply our derived neural networks to long-term well-calibrated UV measurements
9 from TOMS. The TOMS series provides a long-term data record dating back to late 1978
10 (about half a decade before the first ERBE launch) with a few small gaps between that time
11 and the first CERES launch. We should be able to apply NN models derived with
12 CERES/OMI to TOMS, provided that all of the input parameters are available and
13 compatible. In place of actual cloud OCPs that are available from OMI, but not from TOMS,
14 we could use a cloud OCP climatology that was developed from OMI data for use in the
15 TOMS total ozone algorithm. The lower spatial resolution of TOMS is not expected to present
16 any difficulties. This approach can also be extended to the future geostationary missions such
17 as TEMPO, GEMS and Sentinel 4.

19 Acknowledgements

20 This material is based upon work supported by the National Aeronautics and Space
21 Administration issued through the Science Mission Directorate for the Aura Science Team
22 managed by Kenneth Jucks and Richard Eckman. We thank the CERES, OMI, MODIS, and
23 GEOS-DAS data processing teams for providing the data used for this study. We would also
24 like to thanks Norman Loeb and Arlindo da Silva for useful discussion and comments during
25 the preparation of manuscript.

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Table 1. Statistical analysis of the input parameter selection exercise. The correlation coefficient (R), Slope (slope), Bias, and standard deviation of OMI-CERES TOA SW flux for eight different NN models are presented. These numbers correspond to daily inter-comparison between OMI and CERES TOA SW flux. Data from January 2007 is used for this exercise.

Model	Parameters	R	Slope	Bias	STD (W m ⁻²)
a	f_{c_mod}	0.971	0.941	0.051	37.1
b	f_{c_mod} , VZA, RAA	0.979	0.959	0.000	31.4
c	LER, SZA, VZA, RAA	0.979	0.959	-0.030	31.1
d	f_{c_mod} , VZA, RAA, O3	0.980	0.960	-0.009	30.9
e	f_{c_mod} , VZA, RAA, OCP	0.981	0.962	-0.004	30.0
f	f_{c_mod} , VZA, RAA, O3, OCP	0.981	0.963	0.002	29.9
g	f_{c_mod} , VZA, RAA, O3, OCP, PW, Wind	0.982	0.964	0.002	29.2
h	f_{c_mod} , VZA, RAA, O3, OCP, PW, Wind, Chl, LER	0.983	0.967	-0.010	28.3

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1 Table 2. Statistical parameters corresponding daily and monthly inter-comparisons of pixel
 2 and grid levels TOA SW flux data from CERES and OMI.

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	N	R	M	I	BIAS	STD	EE5%	EE10%	EE15%
Pixel	8109323	0.96	0.96	10.5	-1.4	47.7	30	53	69
Daily									
0.5°	1512726	0.97	0.96	11.2	0.94	34.4	37	62	77
1°	529679	0.98	0.96	11.2	1.0	27.9	43	69	83
2°	168181	0.98	0.96	11.0	0.33	23.7	50	76	87
5°	35454	0.99	0.96	9.2	-1.8	20.3	60	84	92
10°	10834	0.99	0.97	7.0	-0.0	14.6	69	90	97
Monthly									
0.5°	108620	0.99	0.98	6.1	1.5	12.9	74	93	97
1°	28849	0.99	0.98	6.9	1.2	11.4	79	94	98
2°	7642	0.99	0.96	9.8	0.25	6.6	94	99	100
5°	1325	0.99	0.96	9.9	-1.8	7.0	95	99	100

4 Note: N – Number of pairs, R- correlation coefficient, M- Slope, I- Intercept, Bias – mean of (OMI-CERES in
 5 Wm^{-2}), STD- standard deviation of (OMI-CERES) in Wm^{-2} , EE –Error Envelope for 5%,10%,15% errors. All
 6 flux values have units of Wm^{-2}

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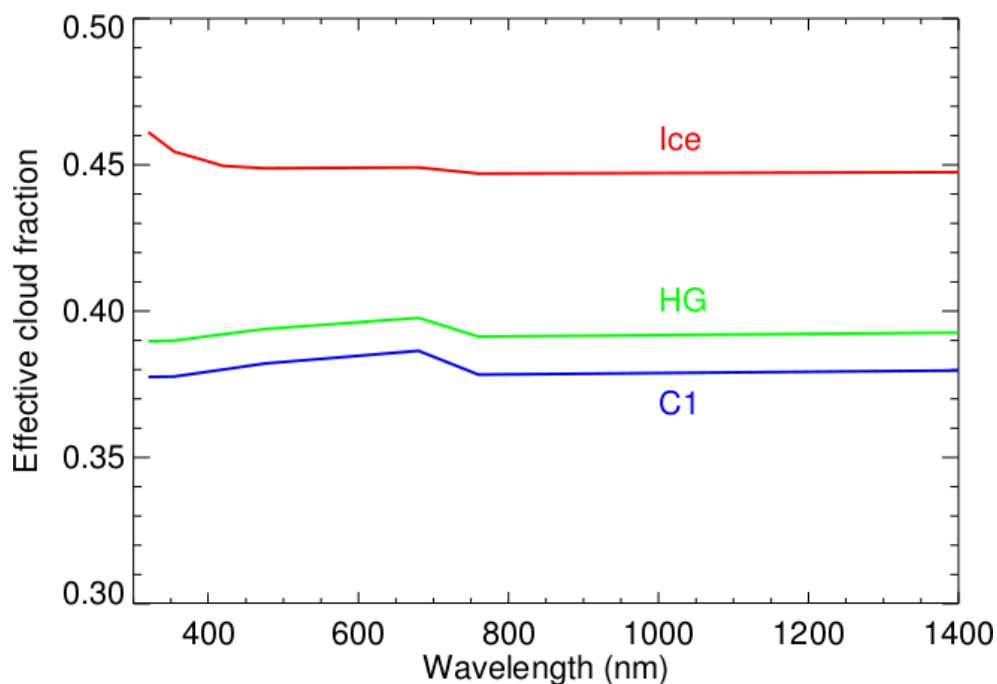
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4 Figure 1. The spectral dependence of the effective cloud fraction for land, $f_g = 0.5$, $SZA = 45^\circ$,
 5 observation at nadir. Red line: ice phase function; Green line: the Heneye-Greenstein
 6 function, Blue line: C1 model.

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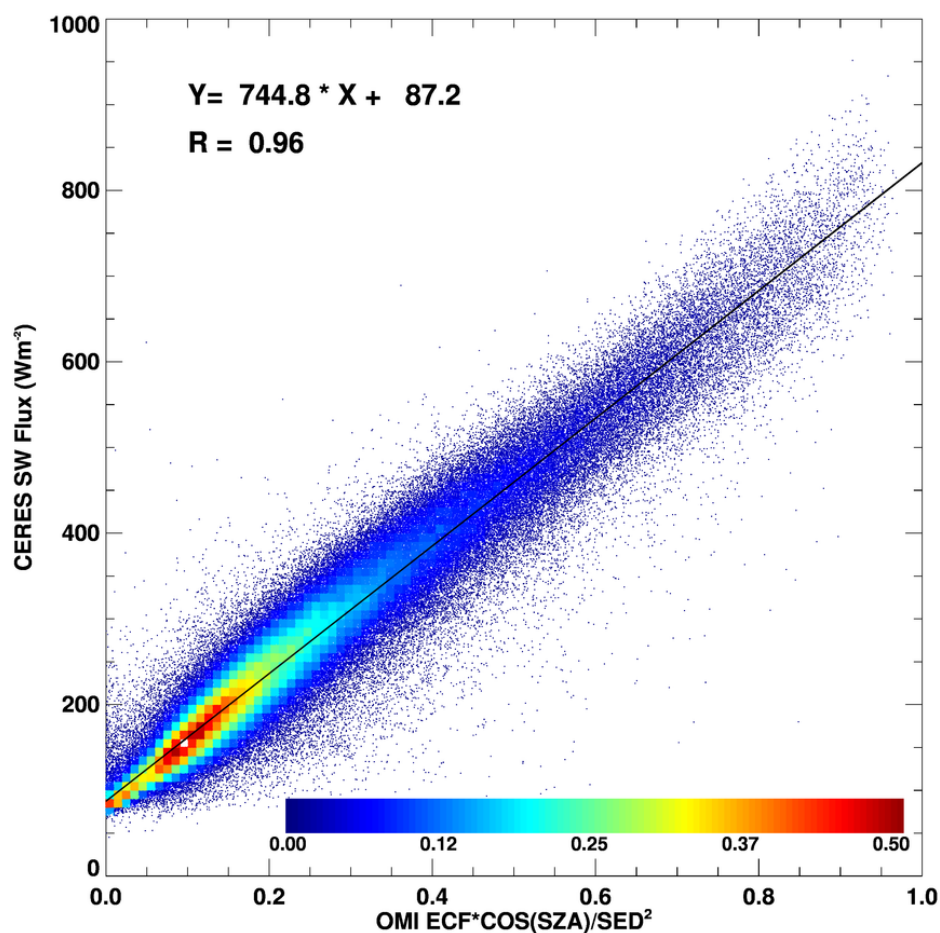
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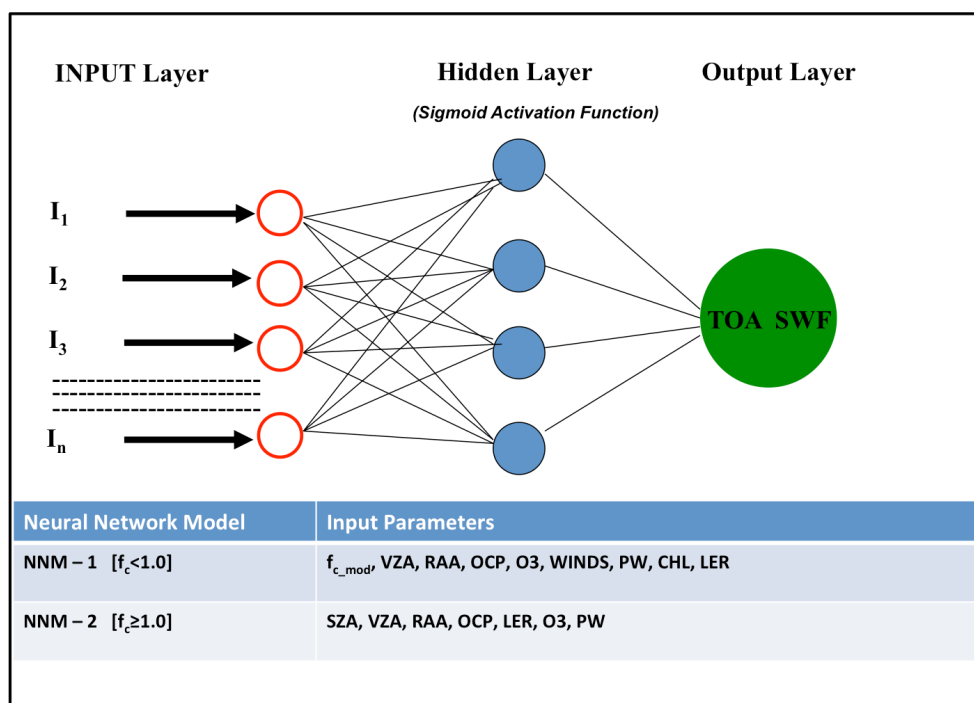
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 3 Figure 2. 2D histogram of effective cloud fraction (ECF or f_c) normalized (i.e. f_{c_mod}) with
 4 respect to incoming solar irradiance and CERES TOA shortwave (SW) flux over ocean.
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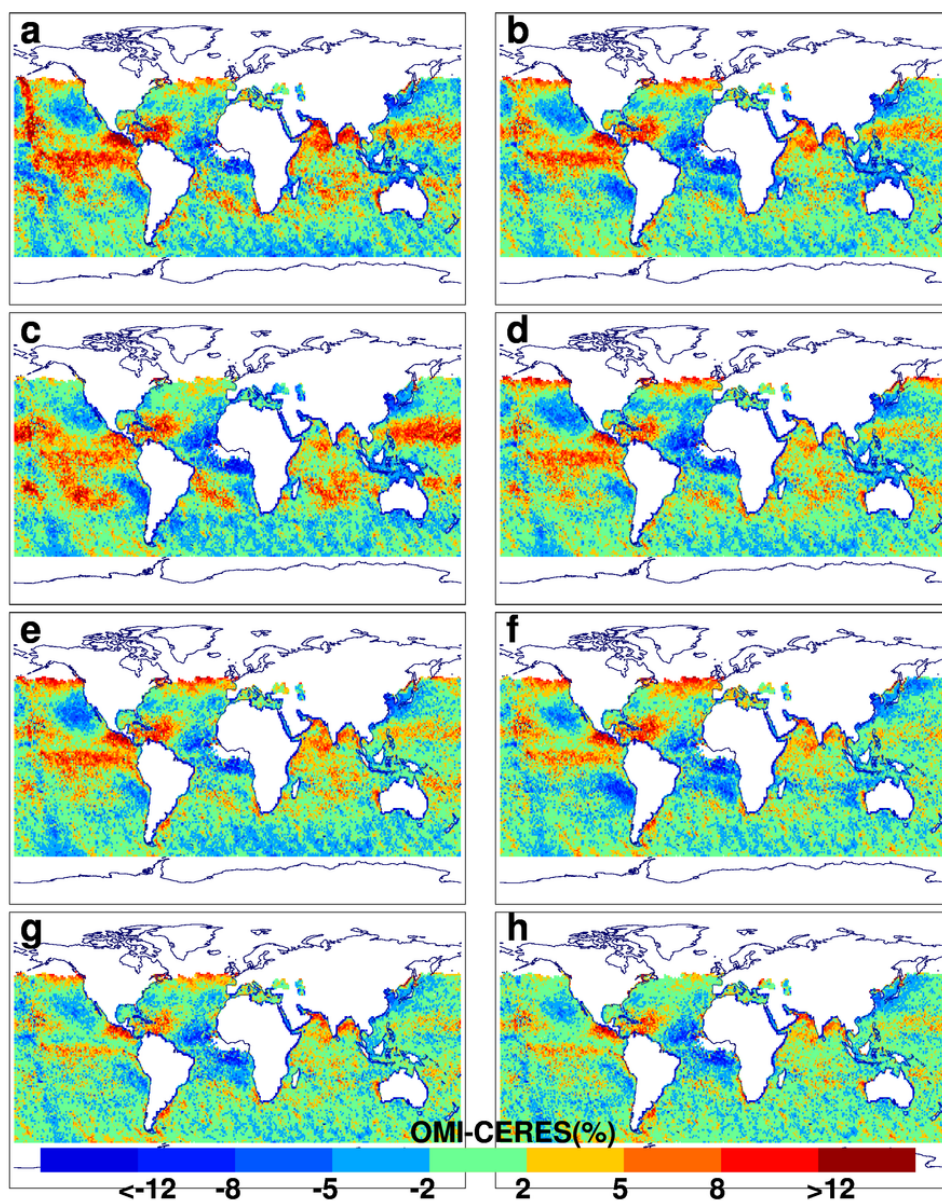
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4 Figure 3. A schematic of the neural network model used for estimation of TOA SW flux with
 5 OMI UV measurements. The table in the bottom lists all the input parameters corresponding
 6 to two NN models used.

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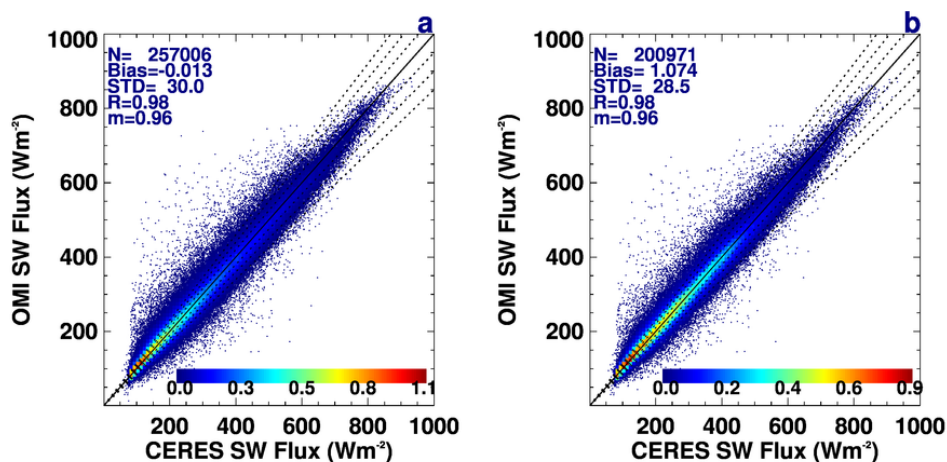


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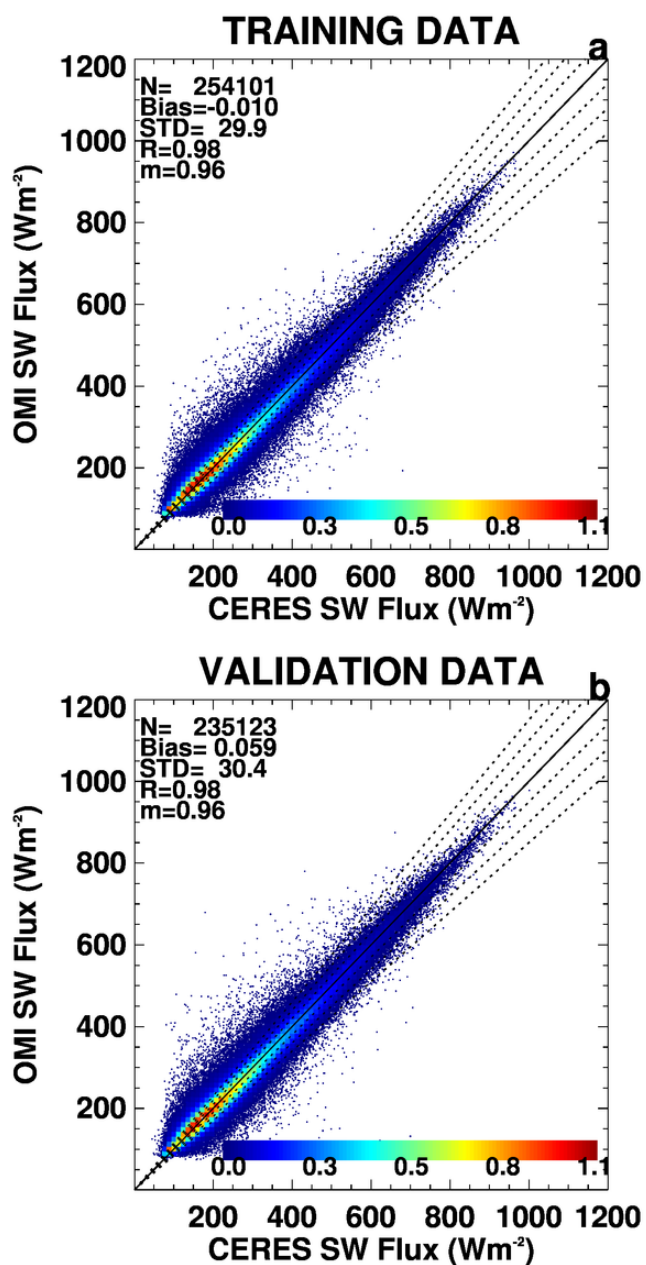
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4 Figure 4. Monthly mean (January 2007) maps of OMI minus CERES TOA SW flux (%) for
 5 eight different NN models. The letters on the map corresponds to model number in the table 1.

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 3 Figure 5. 2D histograms of daily CERES and OMI SW flux for month of January 2007. a) NN
 4 is trained and applied to January 2007 b) NN is trained on January 2007 and applied to
 5 January 2006.
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 3 Figure 6. Similar to Fig. 5 but showing training (top) and validation (bottom) results from two
 4 NN models (input parameters listed in Figure 3, model h for $f_c < 1$ and as in Figure 3 for $f_c > 1$)
 5 as final selected models for estimation of TOA SW Flux.

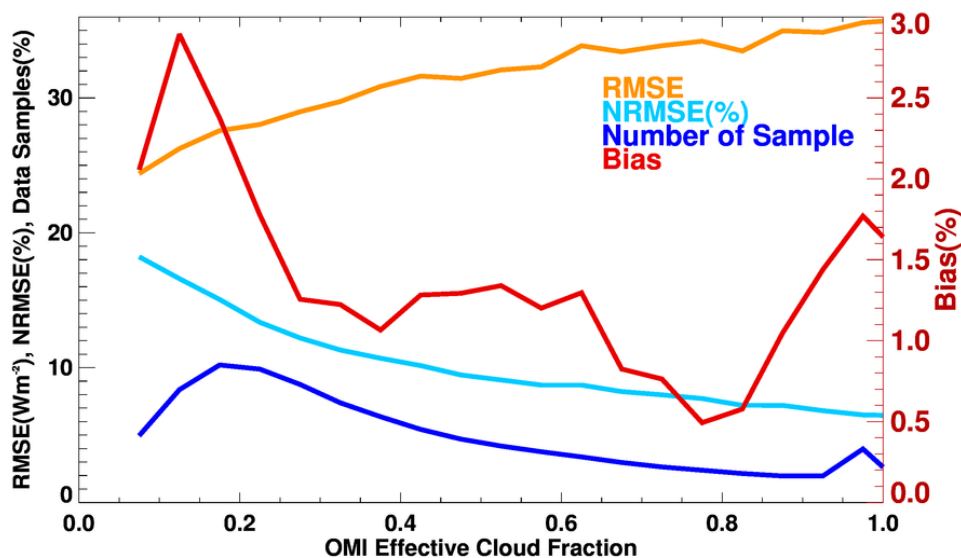
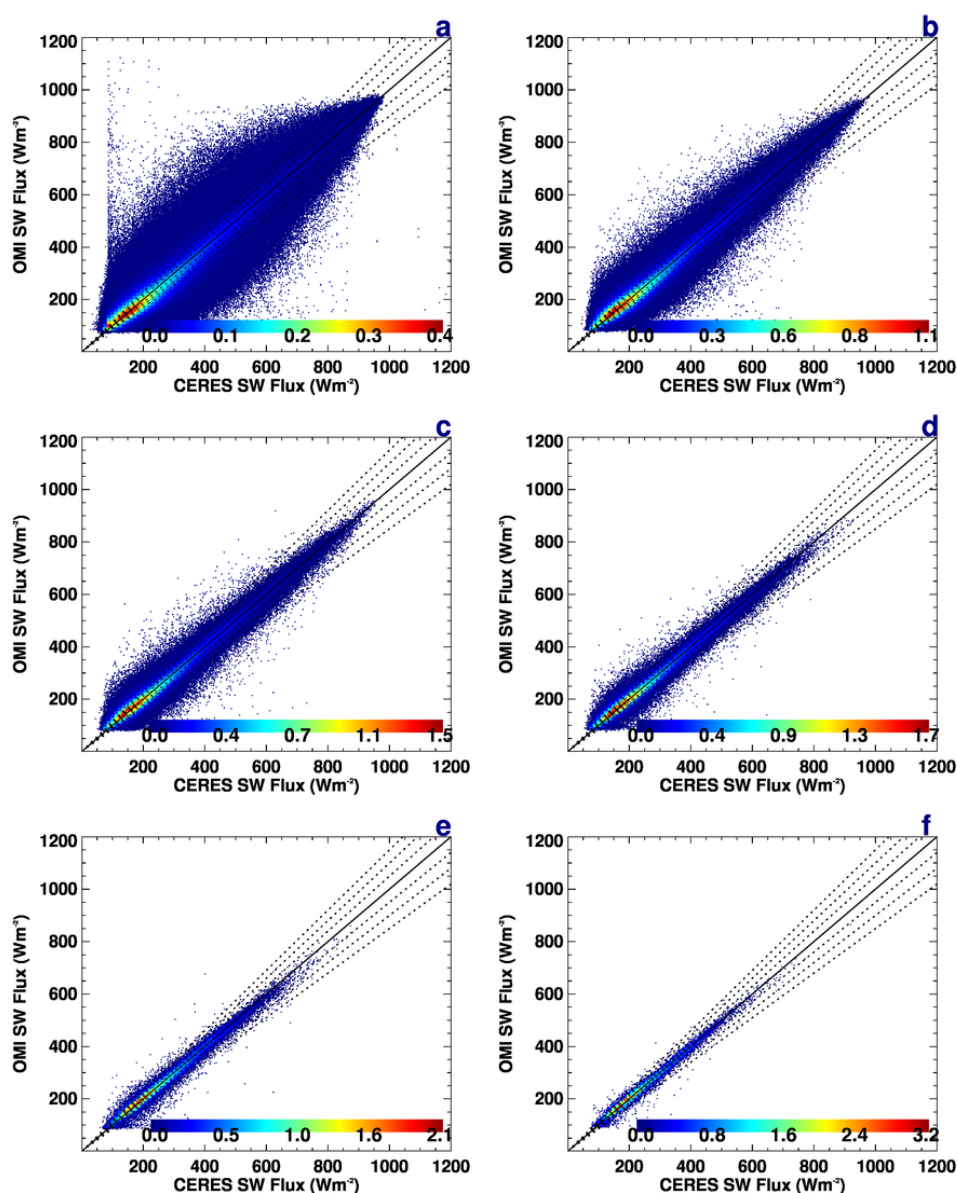
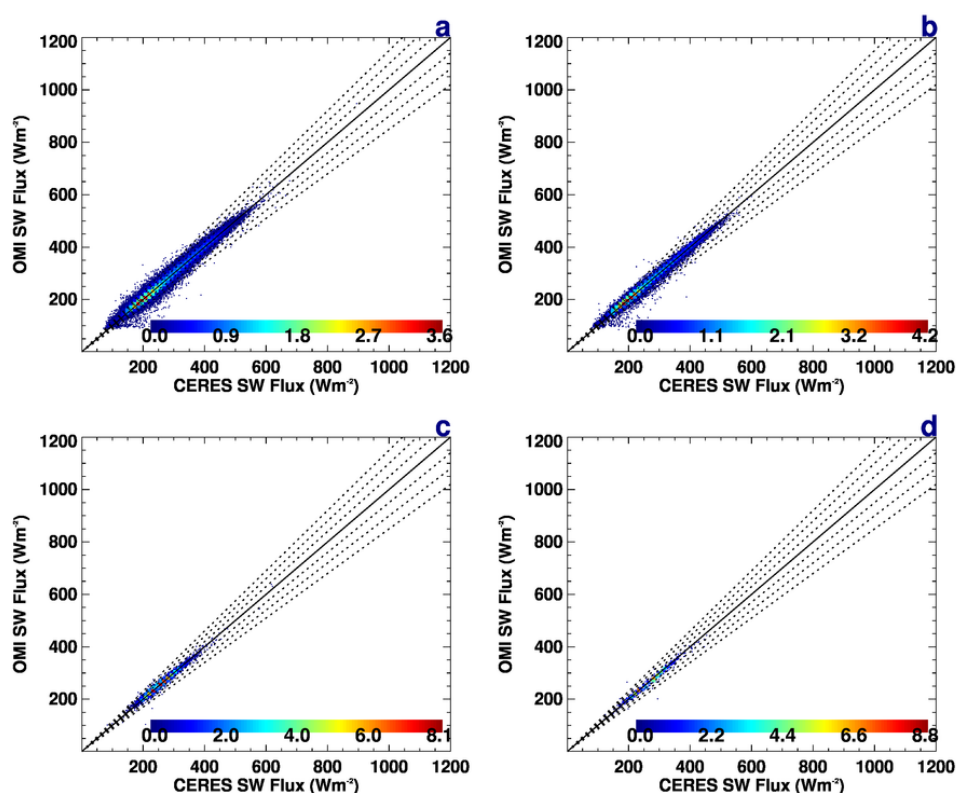


Figure 7. Root mean squared errors (RMSE), normalized RMSE (NRMSE in %), data samples (%), and bias (%) in training and validation data sets as a function of effective cloud fraction for the data presented in Figure 4.



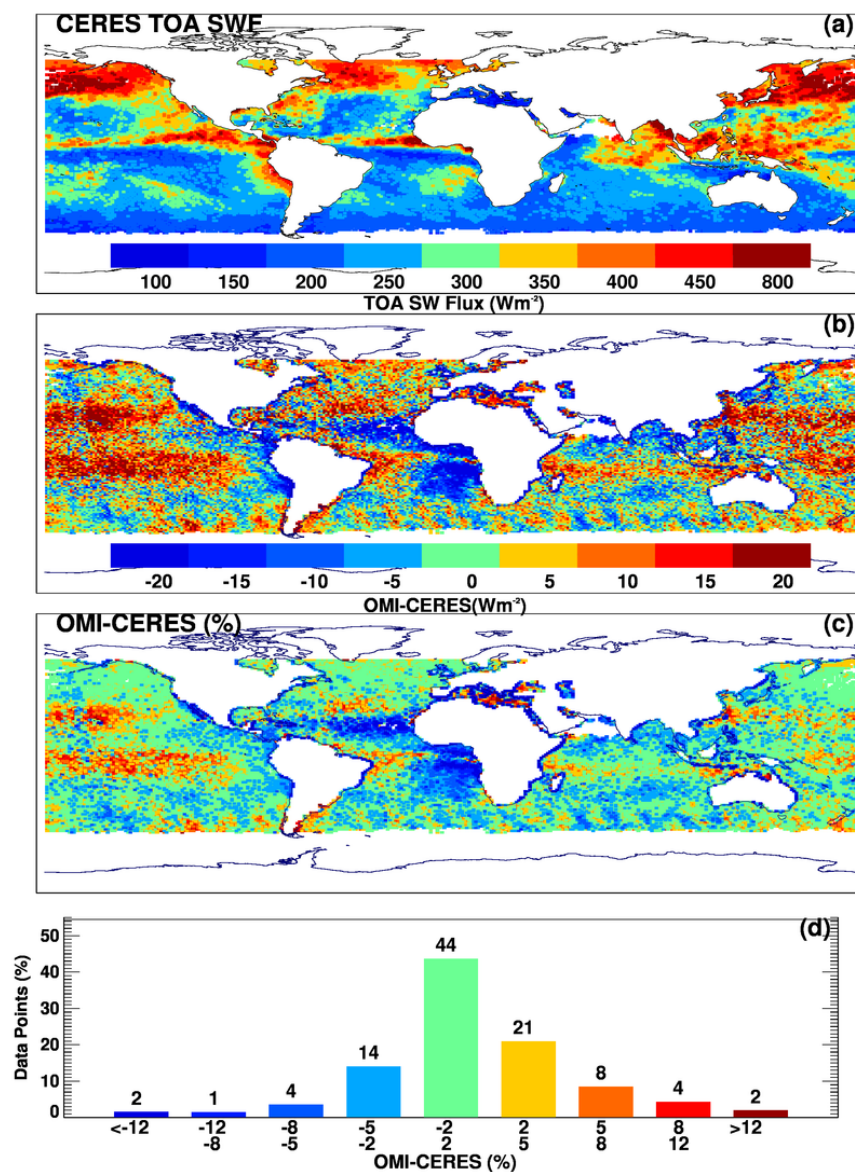
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 3 Figure 8. 2D histograms of the daily OMI and CERES TOA SW Flux averaged over different
 4 spatial grid sizes for the month of July 2007. a) at OMI's native pixel resolution b) 0.5x0.5
 5 degree c) 1x1 degree d) 2x2 degree e) 5x5 degree and f) 10x10 degree. The corresponding
 6 statistical parameters are listed in table 2.

7

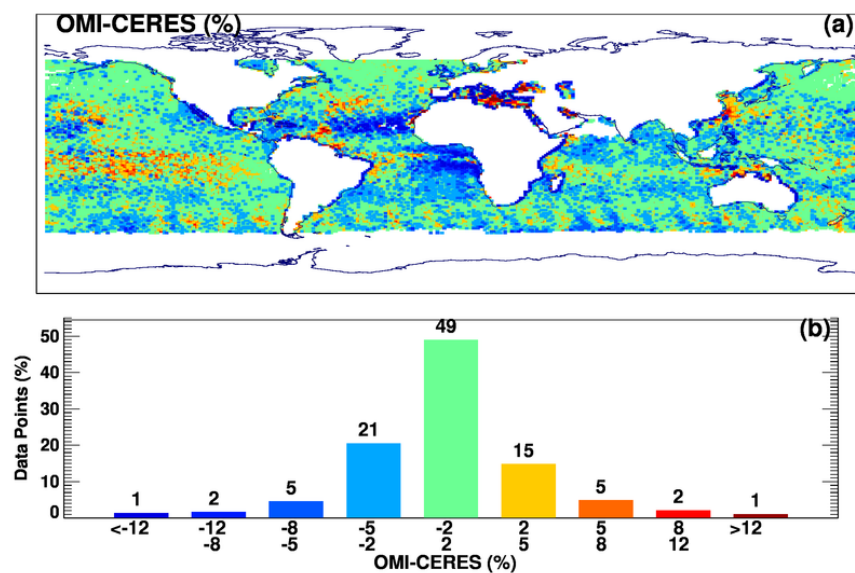


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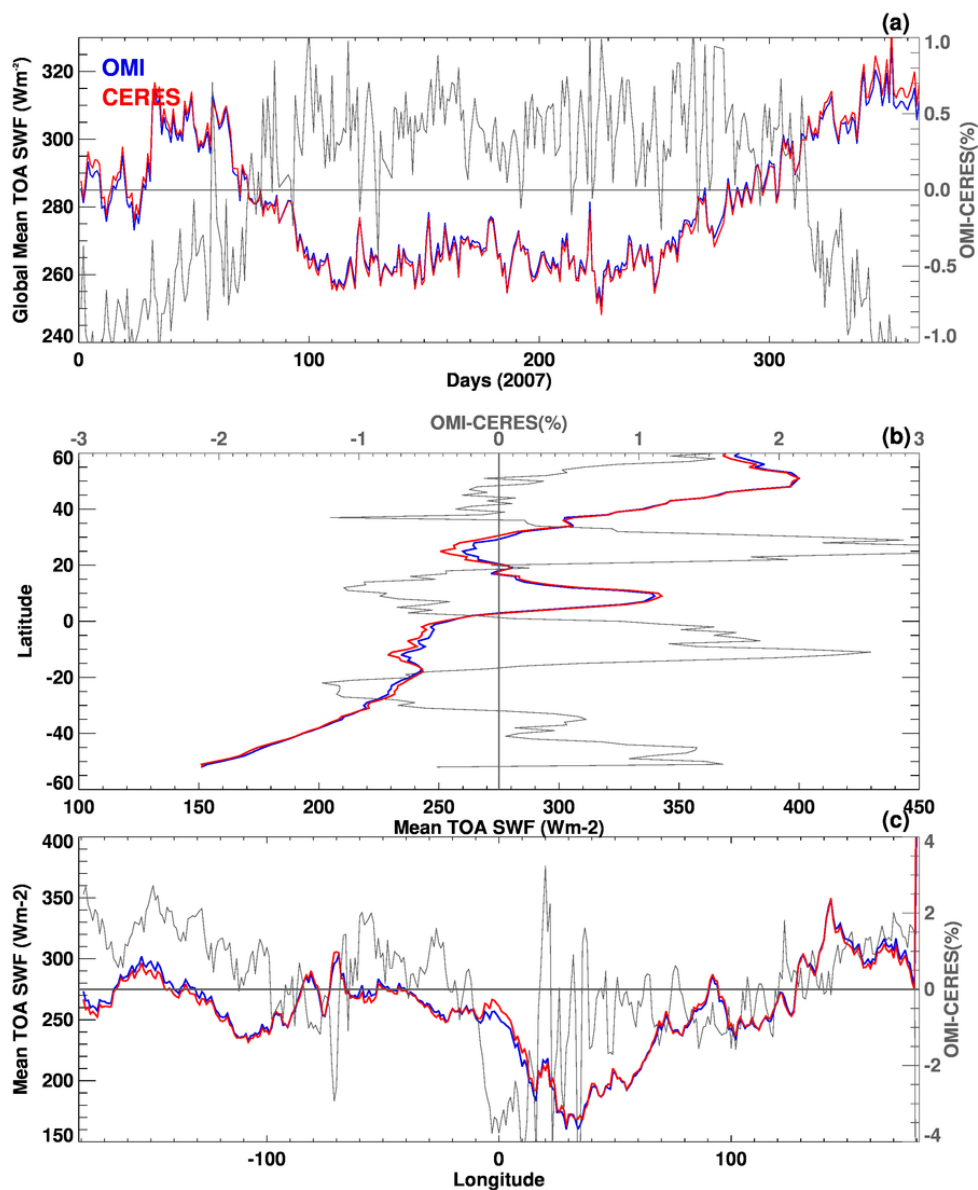
Figure 9. Similar to Fig. 8 but for monthly mean data (July 2007) OMI and CERES TOA SW flux averaged over different spatial grid sizes. a) 0.5x0.5 degree b) 1x1 degree c) 2x2 degree d) 5x5 degree. The corresponding statistical parameters are listed in table 2.



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 2
 3 Figure 10. a) Monthly mean (July 2007) spatial distribution of TOA SW flux from CERES; b)
 4 OMI-CERES (in Wm^{-2}); c) OMI-CERES (%); d) histogram of OMI-CERES % . The colors in
 5 (c) correspond to histogram colors in (d).



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 3 Figure 11. Similar to Fig. 10c and d, but with a NN trained using data from the OMI cloud
 4 O₂-O₂ product.
 5

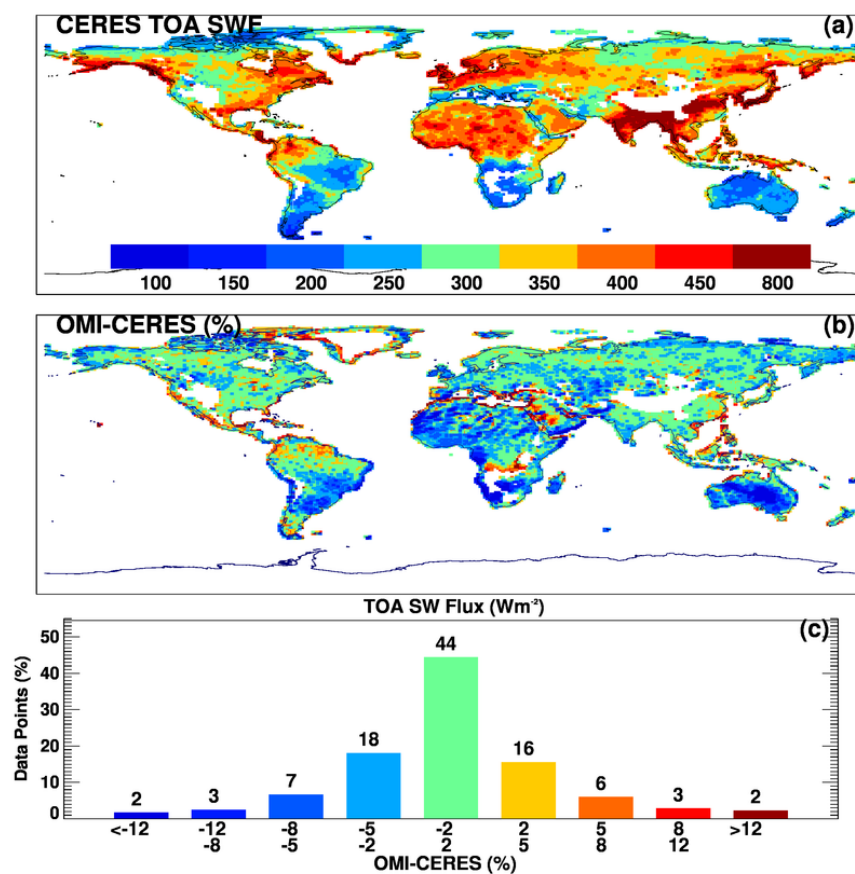


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 3 Figure 12. a) The daily global mean time series of TOA SW fluxes from NN and CERES and
 4 also shows OMI-CERES (%) on secondary y-axis; b) Mean TOA SW flux from CERES and
 5 OMI averaged along each latitude belt, and also shown OMI-CERES (%) on secondary x-
 6 axis. c) same as b but along longitude belts. The data used in b and c are from July 2007.

7



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4 Figure 13. Similar to Fig. 10 (a, c, d) except over land.